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Alternative Description of Rigid Body Kinematics and Quantum Mechanical Angular Momenta

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Two-Axes Decompositions

Given two non-parallel unit vectors $\boldsymbol{\hat{c}}_{1,2} \in \mathbb{S}^2$ the NSC for decomposing

$$\mathcal{R}(\mathbf{n},\phi) = \mathcal{R}(\hat{\mathbf{c}}_2,\phi_2)\mathcal{R}(\hat{\mathbf{c}}_1,\phi_1)$$

is the coincidence of the matrix entries

$$r_{21} = g_{21}$$

and the solution has the form

$$\phi_1 = 2 \arctan \frac{(\mathbf{n}, \, \hat{\mathbf{c}}_1 \times \hat{\mathbf{c}}_2)}{(\mathbf{n}, \, \hat{\mathbf{c}}_{[1})g_{2]1}}, \qquad \phi_2 = 2 \arctan \frac{(\mathbf{n}, \, \hat{\mathbf{c}}_1 \times \hat{\mathbf{c}}_2)}{(\mathbf{n}, \, \hat{\mathbf{c}}_{[2})g_{1]2}}$$

where we denote

$$r_{ij} = (\hat{\mathbf{c}}_i, \mathcal{R}(\mathbf{n}, \phi) \, \hat{\mathbf{c}}_j), \qquad g_{ij} = (\hat{\mathbf{c}}_i, \, \hat{\mathbf{c}}_j), \qquad a_{[i}b_{j]} = a_i b_j - a_j b_i.$$

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Euler Angles and Conjugation

Consider the Euler decomposition of a rotation $\mathcal{R}\in\mathsf{SO}(3)$

$$\mathcal{R}(\mathbf{n},\phi) = \mathcal{R}(\hat{\mathbf{c}}_1,\psi)\mathcal{R}(\hat{\mathbf{c}}_2,\vartheta)\mathcal{R}(\hat{\mathbf{c}}_1,\varphi)$$

where ϕ , φ , ϑ and ψ denote the angles and **n**, $\hat{\mathbf{c}}_{1,2}$ - the invariant axes (with the Davenport condition $\hat{\mathbf{c}}_1 \perp \hat{\mathbf{c}}_2$). A simple conjugation yields

$$\mathcal{R}_{1}(\psi)\mathcal{R}_{2}(\vartheta)\mathcal{R}_{1}(\varphi) = \mathcal{R}_{1}(\psi)\mathcal{R}_{2}(\vartheta)\mathcal{R}_{1}^{-1}(\psi)\mathcal{R}_{1}(\varphi+\psi)$$

so we have an equivalent two-factor decomposition with respect to a (shifted) pair of orthogonal axes

$$\mathcal{R}(\mathbf{n},\phi) = \mathcal{R}\left(\mathcal{R}(\hat{\mathbf{c}}_1,\psi)\,\hat{\mathbf{c}}_2,\,\vartheta\right)\mathcal{R}(\hat{\mathbf{c}}_1,\,\varphi+\psi).$$

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One Particular Solution

It is not difficult to satisfy the condition $\mathit{r}_{21} = \mathit{g}_{21}$ choosing arbitrary $\hat{\mathbf{c}}_1$

$$\hat{\mathbf{c}}_2 = \lambda \, \hat{\mathbf{c}}_1 \! imes \! \mathcal{R}(\mathbf{c}) \, \hat{\mathbf{c}}_1, \qquad \lambda = (1 - r_{11}^2)^{-1/2}$$

and as for the solutions, we have (note that $\phi_1=artheta$ and $\phi_2=arphi+\psi)$

$$\phi_1=2\, {
m arctan}\,
ho_1, \qquad \phi_2={
m arccos}\, r_{11}$$

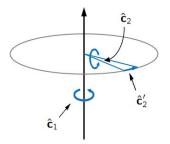
with the notation $\rho_i = \tan \frac{\phi}{2}(\mathbf{n}, \hat{\mathbf{c}}_i)$. One exception is the setting

$$\mathcal{R}(\mathbf{n},\pi) = 2 \, \mathbf{n} \otimes \mathbf{n}^t - \mathcal{I}, \qquad \mathbf{n} \perp \hat{\mathbf{c}}_1$$

that yields a one-parameter set of solutions. Choosing $\hat{\boldsymbol{c}}_2 \perp \hat{\boldsymbol{c}}_1$ one has

$$\mathcal{R}(\mathbf{n},\pi) = \mathcal{R}(\hat{\mathbf{c}}_2,\pi)\mathcal{R}(\hat{\mathbf{c}}_1,\phi_1), \qquad \phi_1 = 2\measuredangle(\mathbf{c}_2,\mathbf{n}).$$

The Orthonormal Frame



One easily constructs a basis with a third vector

$$\hat{\mathbf{c}}_3 = \hat{\mathbf{c}}_1 \times \hat{\mathbf{c}}_2 = \lambda \left[r_{11} \mathcal{I} - \mathcal{R} \right] \hat{\mathbf{c}}_1.$$

In order to parameterize the SO(3) we choose the third coordinate κ as the normal component of the rate, at which $\hat{\mathbf{c}}_2$ varies with \mathcal{R} , i.e.,

$$\hat{\mathbf{c}}_2' = \kappa' \, \hat{\mathbf{c}}_1 \times \, \hat{\mathbf{c}}_2.$$

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Kinematics

The $\{\phi_1,\phi_2,\kappa\}$ coordinates provide the kinematic equations in the form

$$\begin{split} \dot{\phi}_1 &= & \Omega_1 - \Omega_3 \tan \frac{\phi_2}{2} \\ \dot{\phi}_2 &= & \Omega_2 \\ \dot{\kappa} &= & \Omega_1 + \Omega_3 \cot \phi_2 \end{split}$$

where Ω_k denote the components of the angular velocity in the so chosen basis. Inverting the matrix of the above system, one easily obtains

$$\Omega_1 = \dot{w} - \cos v \, \dot{u}$$
$$\Omega_2 = \dot{v}$$
$$\Omega_3 = \sin v \, \dot{u}$$

where we make use of the notation $u = \kappa - \phi_1$, $v = \phi_2$ and $w = \kappa$.

Dynamics

Consider the free Euler equations for a rotational inertial ellipsoid

$$\ddot{u} = -\cot v \, \dot{u} \, \dot{v}$$

$$\ddot{v} = \mu(\cos v \, \dot{u} - \dot{w}) \sin v \, \dot{u}$$

$$\ddot{w} = (\mu \sin v - \csc v) \dot{u} \dot{v}$$

with $I_1 = I_2 = I$ and $\mu = 1 - I_3/I$. One has $\Omega_3 = {\rm const.}$ and

$$v = a\cos(\omega t + \varphi_{\circ}) + b, \qquad \omega = \mu \,\Omega_3.$$

The kinematic equations then yield directly

$$\Omega_1(t) = a\omega \cos(\omega t + \varphi_\circ), \qquad \Omega_2(t) = -a\omega \sin(\omega t + \varphi_\circ)$$

while for the u and v variables one ends up with

$$u = \mp \frac{1}{\mu} \int \frac{\csc v \,\mathrm{d}v}{\sqrt{a^2 - (v - b)^2}} \,, \qquad w = \mp \frac{1}{\mu} \int \frac{\cot v + \mu(v - b)}{\sqrt{a^2 - (v - b)^2}} \,\mathrm{d}v.$$

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Infinitesimal Variations

Infinitesimal left and right deck transformations yield the differential

$$\mathrm{d}\phi = \sin\phi \left(\csc\phi_1 \,\mathrm{d}\phi_1 + \frac{1 - \cos\phi_2}{\cos\phi_1 - \cos\phi} \,\csc\phi_2 \,\mathrm{d}\phi_2 + \csc\phi_1 \csc\phi_2 \,\mathrm{d}\kappa \right)$$

as well as the components of the angular momentum operator

$$L_{1} = \frac{\partial}{\partial\phi_{1}}$$

$$L_{2} = \sin\phi_{1}\tan\frac{\phi_{2}}{2}\frac{\partial}{\partial\phi_{1}} + \cos\phi_{1}\frac{\partial}{\partial\phi_{2}} + \sin\phi_{1}\csc\phi_{2}\frac{\partial}{\partial\kappa}$$

$$L_{3} = \cos\phi_{1}\tan\frac{\phi_{2}}{2}\frac{\partial}{\partial\phi_{1}} - \sin\phi_{1}\frac{\partial}{\partial\phi_{2}} + \cos\phi_{1}\csc\phi_{2}\frac{\partial}{\partial\kappa}$$

The Laplacian

The associated Laplace operator (quantum Hamiltonian) has the form

$$\Delta = \sec^2 \frac{\phi_2}{2} \frac{\partial^2}{\partial \phi_1^2} + \frac{\partial^2}{\partial \phi_2^2} - \tan \frac{\phi_2}{2} \frac{\partial}{\partial \phi_2} + \csc^2 \phi_2 \frac{\partial^2}{\partial \kappa^2}$$

Using the notation $(\phi_1,\phi_2)
ightarrow (lpha,artheta)$ one may rewrite the above as

$$\Delta = \sec^2 \frac{\vartheta}{2} \left[\frac{\partial^2}{\partial \alpha^2} + \frac{\partial}{\partial \vartheta} \left(\cos^2 \frac{\vartheta}{2} \frac{\partial}{\partial \vartheta} \right) \right] + \csc^2 \vartheta \frac{\partial^2}{\partial \kappa^2}.$$

For $\kappa = \text{const.}$ we obtain the restriction on the quadric $r_{21} = g_{21}$

$$\Delta_0 = \sec^2 \frac{\vartheta}{2} \frac{\partial^2}{\partial \alpha^2} + \frac{\partial^2}{\partial \vartheta^2} - \tan \frac{\vartheta}{2} \frac{\partial}{\partial \vartheta}.$$

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The Hyperbolic Case

The factors $\epsilon_k = \hat{\mathbf{c}}_k^2$ distinguish between space-like ($\epsilon_k = 1$), time-like ($\epsilon_k = -1$) and light-like ($\epsilon_k = 0$) vectors in $\mathbb{R}^{2,1}$. We may denote

$$au= au = au + rac{\phi}{2} \ (\epsilon \!=\! 1), \qquad au = au = au rac{\phi}{2} \ (\epsilon \!=\! -1), \qquad au = rac{\phi}{2} \ (\epsilon \!=\! 0)$$

thus unifying angles and rapidities. Note also the isotropic singularity

Theorem

If $\hat{\mathbf{c}}_{1,2} \in \mathbf{c}_{\circ}^{\perp}$ with $\mathbf{c}_{\circ}^2 = 0$, we may decompose a pseudo-rotation

$$\Lambda(\mathbf{n},\tau) = \Lambda(\hat{\mathbf{c}}_2,\tau_2)\Lambda(\hat{\mathbf{c}}_1,\tau_1), \qquad \Lambda \in \mathrm{SO}(2,1)$$

if and only if $\mathbf{n} \in \mathbf{c}_{\circ}^{\perp}$ lies the same tangent plane to the null cone.

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The Monodromy Matrix

The quantum mechanical monodromy matrix

$$\mathcal{M} = \left(egin{array}{cc} 1/ar{t} & -ar{r}/ar{t} \ -r/t & 1/t \end{array}
ight) \in {\sf SU}(1,1)$$

relates left and right free particle asymptotic solutions

$$\begin{split} \Psi(k,x) &\sim \mathrm{e}^{\mathrm{i}kx} + r(k) \,\mathrm{e}^{-\mathrm{i}kx}, \qquad x o -\infty \ \Psi(k,x) &\sim t(k) \,\mathrm{e}^{\mathrm{i}kx}, \qquad x o \infty. \end{split}$$

In a standard split-quaternion basis $\ensuremath{\mathcal{M}}$ may be decomposed as

$$\mathcal{M} \rightarrow \zeta_{\mathcal{M}} = (\Re(t), -\Re(r\overline{t}), \Im(r\overline{t}), \Im(t))^{t}$$

associating $t \in \mathbb{R}$ with pure boosts and r = 0 with pure rotations.

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Decomposition of Scattering Potentials

Choosing $\hat{\boldsymbol{c}}_1$ to be aligned with the z-axis we find

$$\phi_1 = 2\theta, \qquad \theta = \arg(t)$$

as well as

$$\tau_2 = \sqrt{1 - |t|^2} \quad \Rightarrow \quad \phi_2 = 2 \operatorname{arccosh} |t|^{-1}$$

which finally yields

$$\mathcal{M} = rac{1}{|t|} \left(egin{array}{cc} 1 & -\overline{r}\mathrm{e}^{\,2\mathrm{i} heta} \ -r\mathrm{e}^{-2\mathrm{i} heta} & 1 \end{array}
ight) \left(egin{array}{cc} \mathrm{e}^{\mathrm{i} heta} & 0 \ 0 & \mathrm{e}^{-\mathrm{i} heta} \end{array}
ight)$$

and we decompose the monodromy into a product of a pure phase shift and a phase preserving scattering.

Extension to SO(3, 1)

The local isomorphism

$$\mathrm{SO}^+(3,1)\cong\mathrm{SO}(3,\mathbb{C})$$

allows for extending via complexification e.g using biquaternions.

The existence of invariant axes is ensured by the Plücker relations

$$(\Re \, au \, {f n}, \Im \, au \, {f n}) = (\Re \, au_k \hat{f c}_k, \Im \, au_k \hat{f c}_k) = 0$$

and decomposability with real scalar parameters demands

$$(\Re \, \tau \, \mathbf{n}, \Im \, \tau_k \hat{\mathbf{c}}_k) + (\Re \, \tau_k \hat{\mathbf{c}}_k, \Im \, \tau \, \mathbf{n}) = 0$$

which projects the problem to a three-dimensional hyperplane.

Similar arguments (and Plücker relations) hold for the groups

$$SO(4)$$
, $SO(2,2)$, $SO^{*}(4)$.

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Thomas Precession and Wigner Little Groups

Consider the generalized Rodrigues' vector for SO(3, 1)

$$\mathbf{c} = \tau \mathbf{n} = \boldsymbol{\alpha} + \mathrm{i}\beta \in \mathbb{CP}^3.$$

In the Plücker setting $\pmb{\alpha} \perp \pmb{\beta}$ we may write

$$\Lambda(\alpha + \mathrm{i}\beta) = \Lambda(\mathrm{i}\tilde{\beta}_{+})\Lambda(\alpha) = \Lambda(\alpha)\Lambda(\mathrm{i}\tilde{\beta}_{-}), \qquad \tilde{\beta}_{\pm} = \frac{\mathcal{I} \pm \alpha^{\times}}{1 + \alpha^{2}}\beta.$$

One example is the Thomas precession, in which

$$oldsymbol{lpha} = rac{eta_1 imes eta_2}{1 + eta_1 \cdot eta_2}, \qquad eta = rac{eta_1 + eta_2}{1 + eta_1 \cdot eta_2}$$

where $i\beta_{1,2}$ are the Rodrigues' vectors of the two consecutive boosts.

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A Numerical Example

Consider a proper Lorentz transformation

$$\Lambda = \frac{1}{5} \left(\begin{array}{rrrr} -6 & 2 & -1 & -4 \\ -4 & -3 & -6 & -6 \\ -3 & -4 & 2 & -2 \\ 6 & 2 & 4 & 9 \end{array} \right)$$

and let us choose $\hat{\bm{c}}_1=(1{+}\mathrm{i},\,1{-}\mathrm{i},\,1)$, so that $\Lambda(\hat{\bm{c}}_1,\tau_1)$ preserves

$$\boldsymbol{\varsigma} = (1, 1, 1, 0)^t.$$

Then, we easily obtain $\Lambda = \Lambda_2 \Lambda_1$ with

$$\Lambda_1 = \frac{1}{17} \begin{pmatrix} \begin{array}{cccc} -15 & 8 & 24 & 24 \\ -8 & -15 & 40 & 40 \\ 40 & 24 & -47 & -64 \\ -40 & -24 & 64 & 81 \end{array} \end{pmatrix}, \quad \Lambda_2 = \frac{1}{85} \begin{pmatrix} \begin{array}{cccc} 178 & 138 & -401 & -452 \\ 36 & 77 & -334 & -334 \\ 109 & 224 & -334 & -334 \\ 109 & 224 & -438 & -506 \\ -194 & -278 & 676 & 761 \end{array} \end{pmatrix}$$

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Rational Coordinates

Given a rotation matrix $\mathcal{R} \in SO(3)$ with rational coefficients and a rational unit vector $\hat{\mathbf{c}}_1$ we may construct $\hat{\mathbf{c}}_2 = \hat{\mathbf{c}}_1 \times \mathcal{R}\hat{\mathbf{c}}$, which yields

$$\tau_1 = \rho_1, \qquad \tau_2 = \frac{1}{1 + r_{11}}$$

in the Euclidean case and

$$au_1 = \epsilon_1^{-1} \rho_1, \qquad au_2 = (\epsilon_1 + r_{11})^{-1}$$

for SO(3, 1), so the two factors are rational as well.

Recommended Readings

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Thank You!



THANKS FOR YOUR PATIENCE!