### Geodesic Mappings and Einstein Spaces

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### Topics

- **(**) Geodesic mapping theory for  $V_n \to \bar{V}_n$  of class  $C^1$
- 2 Geodesic mapping theory for  $V_n \rightarrow \bar{V}_n$  of class  $C^2$
- **③** Geodesic mapping between  $V_n \in C^r$  (r > 2) and  $\bar{V}_n \in C^1$
- On geodesic mappings of Einstein spaces

# 1. Geodesic mapping theory for $V_n \to \bar{V}_n$ of class $C^1$

Assume the (pseudo-) Riemannian manifolds  $V_n = (M, g, \nabla)$  and  $\bar{V}_n = (\bar{M}, \bar{g}, \bar{\nabla})$ . Here  $V_n$ ,  $\bar{V}_n \in C^1$ , i.e.  $g, \bar{g} \in C^1$  which means that their components  $g_{ij}, \bar{g}_{ij} \in C^1$ .

### Definition

A diffeomorphism  $f: V_n \to \overline{V}_n$  is called a *geodesic mapping* of  $V_n$  onto  $\overline{V}_n$  if f maps any geodesic in  $V_n$  onto a geodesic in  $\overline{V}_n$ .

A manifold  $V_n$  admits a geodesic mapping onto  $\overline{V}_n$  if and only if the *Levi-Civita equations* 

(1) 
$$\overline{\nabla}_X Y = \nabla_X Y + \psi(X)Y + \psi(Y)X$$

hold for any tangent fields X, Y and where  $\psi$  is a differential form. If  $\psi \equiv 0$  than f is affine or trivially geodesic. In local form:  $\overline{\Gamma}_{ij}^{h} = \Gamma_{ij}^{h} + \psi_i \delta_j^{h} + \psi_j \delta_i^{h}$ , where  $\Gamma_{ij}^{h}(\overline{\Gamma}_{ij}^{h})$  are the Christoffel symbols of  $V_n$  and  $\overline{V}_n$ ,  $\psi_i$  are components of  $\psi$  and  $\delta_i^{h}$  is the Kronecker delta. Equations (1) are equivalent to the following equations

(2) 
$$\bar{g}_{ij,k} = 2\psi_k \bar{g}_{ij} + \psi_i \bar{g}_{jk} + \psi_j \bar{g}_{ik}$$

where "," denotes the covariant derivative on  $V_n$ . It is known that

$$\psi_i = \partial_i \Psi, \quad \Psi = rac{1}{2(n+1)} \ln \left| rac{\det \bar{g}}{\det g} \right|, \quad \partial_i = \partial / \partial x^i.$$

Sinyukov proved that the Levi-Civita equations are equivalent to

(3) 
$$a_{ij,k} = \lambda_i g_{jk} + \lambda_j g_{ik},$$

where

(4) (a) 
$$a_{ij} = e^{2\Psi} \bar{g}^{\alpha\beta} g_{\alpha i} g_{\beta j};$$
 (b)  $\lambda_i = -e^{2\Psi} \bar{g}^{\alpha\beta} g_{\beta i} \psi_{\alpha}.$   
From (3) follows  $\lambda_i = \partial_i \lambda = \partial_i (\frac{1}{2} a_{\alpha\beta} g^{\alpha\beta}).$  On the other hand

(5) 
$$\bar{g}_{ij} = e^{2\Psi}\tilde{g}_{ij}, \quad \Psi = \frac{1}{2}\ln\left|\frac{\det g}{\det g}\right|, \quad \|\tilde{g}_{ij}\| = \|g^{i\alpha}g^{j\beta}a_{\alpha\beta}\|^{-1}.$$

The above formulas are the criterion for geodesic mappings  $V_n \rightarrow \bar{V}_n$  globally as well as locally.

Let  $V_n$  and  $\bar{V}_n \in C^2$ , then for geodesic mappings  $V_n \to \bar{V}_n$  the Riemann and the Ricci tensors transform in this way (6) (a)  $\bar{R}^h_{ijk} = R^h_{ijk} + \delta^h_k \psi_{ij} - \delta^h_j \psi_{ik}$ ; (b)  $\bar{R}_{ij} = R_{ij} - (n-1)\psi_{ij}$ , where  $\psi_{ij} = \psi_{i,j} - \psi_i \psi_j$ ,

and the Weyl tensor of projective curvature, which is defined in the following form

$$W_{ijk}^{h} = R_{ijk}^{h} + \frac{1}{n-1} \left( \delta_{k}^{h} R_{ij} - \delta_{j}^{h} R_{ik} \right),$$

is invariant.

### The integrability conditions of the Sinyukov equations (3)

have the following form

(7) 
$$a_{i\alpha}R^{\alpha}_{jkl} + a_{j\alpha}R^{\alpha}_{ikl} = g_{ik}\lambda_{j,l} + g_{jk}\lambda_{i,l} - g_{il}\lambda_{j,k} - g_{jl}\lambda_{i,k}.$$

After contraction with  $g^{jk}$  we get

(8) 
$$n\lambda_{i,l} = \mu g_{il} - a_{i\alpha} R_l^{\alpha} + a_{\alpha\beta} R_{il}^{\alpha\beta}$$

where  $R^{\alpha}{}_{il}{}^{\beta} = g^{\beta k} R^{\alpha}{}_{ilk}$ ;  $R^{\alpha}_{l} = g^{\alpha j} R_{jl}$  and  $\mu = \lambda_{\alpha,\beta} g^{\alpha\beta}$ .

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3. Geodesic mapping between  $V_n \in C^r$  (r > 2) and  $\overline{V}_n \in C^1$ 

### Theorem 1

If  $V_n \in C^r$  (r > 2) admits geodesic mappings onto  $\bar{V}_n \in C^1$ , then  $\bar{V}_n \in C^r$ .

This Theorem is more strong than following theorem

Theorem 2

If  $V_n \in C^r$  (r > 2) admits geodesic mappings onto  $\bar{V}_n \in C^2$ , then  $\bar{V}_n \in C^r$ .

#### Lemma 1

Let  $\lambda^h \in C^1$  be a vector field and  $\varrho$  a function. If  $\partial_i \lambda^h - \varrho \, \delta^h_i \in C^1$  then  $\lambda^h \in C^2$  and  $\varrho \in C^1$ .

## Sketch of the proof:

The condition  $\partial_i \lambda^h - \varrho \, \delta^h_i \in C^1$  can be written in the following form

(9) 
$$\partial_i \lambda^h - \varrho \delta^h_i = f^h_i(x),$$

where  $f_i^h(x)$  are functions of class  $C^1$ . Evidently,  $\varrho \in C^0$ . For fixed but arbitrary indices  $h \neq i$  we integrate (9) with respect to  $dx^i$ :

$$\lambda^h = \Lambda^h + \int_{x_o^i}^{x^i} f_i^h(x^1, \dots, x^{i-1}, t, x^{i+1}, \dots, x^n) dt,$$

where  $\Lambda^h$  is a function, which does not depend on  $x^i$ . Because of the existence of the partial derivatives of the functions  $\lambda^h$  and the above integrals, also the derivatives  $\partial_h \Lambda^h$  exist. Then we can write (9) for h = i:

(10) 
$$\varrho = -f_h^h + \partial_h \Lambda^h + \int_{x_o^i}^{x^i} \partial_h f_i^h(x^1, \dots, x^{i-1}, t, x^{i+1}, \dots, x^n) dt.$$

Because the derivative with respect to  $x^i$  of the right-hand side of (10) exists, the derivative of the function  $\rho$  exists, too. Obviously  $\partial_i \rho = \partial_h f_i^h - \partial_i f_h^h$ , therefore  $\rho \in C^1$  and from (9) follows  $\lambda^h \in C^2$ .

In a similar way we can prove the following: if  $\lambda^h \in C^r$   $(r \ge 1)$  and  $\partial_i \lambda^h - \varrho \delta^h_i \in C^r$  then  $\lambda^h \in C^{r+1}$  and  $\varrho \in C^r$ .

### Lemma 2

If  $V_n \in C^3$  admits a geodesic mapping onto  $\bar{V}_n \in C^2$ , then  $\bar{V}_n \in C^3$ .

### Skach of the proof

In this case Sinyukov's equations (3) and (8) hold. According to the assumptions  $g_{ii} \in C^3$  and  $\bar{g}_{ii} \in C^2$ . By a simple check-up we find  $\Psi \in C^2$ ,  $\psi_i \in C^1$ ,  $a_{ii} \in C^2$ ,  $\lambda_i \in C^1$  and  $R_{iik}^h, R_{ii}^h, R_{ii}, R_i^h \in C^1.$ From the above-mentioned conditions we easily convince ourselves that we can write equation (8) in the form (9), where  $\lambda^h = g^{h\alpha} \lambda_{\alpha} \in C^1$ ,  $\rho = \mu/n$  and  $f_i^h = (-\lambda^{\alpha} \Gamma^h_{\alpha i} - g^{h\gamma} a_{\alpha\gamma} R_i^{\alpha} + g^{h\gamma} a_{\alpha\beta} R^{\alpha}{}_{i\gamma}{}^{\beta})/n \in C^1.$ From Lemma 1 follows that  $\lambda^h \in C^2$ ,  $\rho \in C^1$ , and evidently  $\lambda_i \in C^2$ . Differentiating (3) twice we convince ourselves that  $a_{ii} \in C^3$ . From this and formula (5) follows that also  $\Psi \in C^3$  and  $\bar{g}_{ii} \in C^3$ .

Further we notice that for geodesic mappings between  $V_n$  and  $\bar{V}_n$  of class  $C^3$  holds the third set of Sinyukov equations:

(11) 
$$(n-1)\mu_{,k} = 2(n+1)\lambda_{\alpha}R_{k}^{\alpha} + a_{\alpha\beta}(2R_{k}^{\alpha}, \beta - R_{k}^{\alpha\beta}, k).$$

If  $V_n \in C^r$  and  $\overline{V}_n \in C^2$ , then by Lemma 2,  $\overline{V}_n \in C^3$  and (11) hold. Because Sinyukov's system (3), (8) and (11) is closed, we can differentiate equations (3) (r-1) times. So we convince ourselves that  $a_{ij} \in C^r$ , and also  $\overline{g}_{ij} \in C^r$  ( $\equiv \overline{V}_n \in C^r$ ).

#### Remark

Because for holomorphically projective mappings of Kähler (and also hyperbolic and parabolic Kähler) spaces hold equations analogical to (3) and (8), from Lemma 1 follows an analog to Theorem 1 for these mappings.

## 4.On geodesic mappings of Einstein spaces

Einstein spaces  $V_n$  are characterized by the condition

 $Ric = const \cdot g$ ,

so  $V_n \in C^2$  would be sufficient.

We remark that spaces of constant curvature are Einstein spaces and Einstein spaces  $V_3$  are always have constant curvature. Therefore many properties of Einstein spaces appear when

$$V \in C^3$$
 and  $n > 3$ .

Moreover, it is known (D.M. DeTurck and J.L. Kazdan) that an Einstein space  $V_n$  belongs to  $C^{\omega}$ , i.e., for all points of  $V_n$ , there exists local coordinate system x for which  $g_{ij}(x) \in C^{\omega}$  (analytic coordinate system).

It is known that Riemannian spaces of constant curvature form a closed class with respect to geodesic mappings (Beltrami theorem).

### Theorem 3

If the Einstein space  $V_n$  admits a nontrivial geodesic mapping onto a (pseudo-) Riemannian space  $\bar{V}_n$ , then  $\bar{V}_n$  is an Einstein space. In 1978 in the PhD thesis Mikeš proved that above Theorem holds locally for  $V_n \in C^3$  and  $\bar{V}_n \in C^3$ . From Theorem 2 this Theorem holds for  $V_n \in C^3$  and  $\bar{V}_n \in C^1$ . Moreover from results by DeTurck this Theorem holds GLOBALLY and exists common coordinate system in which  $V_n \in C^{\omega}$  and  $\bar{V}_n \in C^{\omega}$ .

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Thank you for your attention!